

Name: \_\_\_\_\_

## Intro Astro Lab Prep Quiz: Lab 10 Stellar Spectra

**Instructions:** There are 10 multiple-choice problems each worth 1 marks for a total of 10 marks altogether. Choose the **BEST** answer, completion, etc., and **DARKEN** fully the appropriate circle on the table provided below. Read all responses carefully. **NOTE** long detailed responses won't depend on hidden keywords: keywords in such responses are bold-faced capitalized.

This is a 10 minute quiz.

### Answer Table for the Multiple-Choice Questions

	a	b	c	d	e		a	b	c	d	e
1.	O	O	O	O	O	6.	O	O	O	O	O
2.	O	O	O	O	O	7.	O	O	O	O	O
3.	O	O	O	O	O	8.	O	O	O	O	O
4.	O	O	O	O	O	9.	O	O	O	O	O
5.	O	O	O	O	O	10.	O	O	O	O	O

1. "Let's play *Jeopardy!* For \$100, the answer is: It is the range of all possible wavelengths of electromagnetic radiation. At least as an ideal limit, the wavelengths form a continuum (like real numbers) ranging from arbitrarily close to zero to arbitrarily close to infinity. Real processes may limit the actual range of wavelengths, but we really don't know where those limits are."

What is \_\_\_\_\_, Alex?

- a) white light    b) white noise    c) colored light    d) the energy spectrum  
e) the electromagnetic spectrum

2. "Let's play *Jeopardy!* For \$100, the answer is: Because the Earth's atmosphere is very transparent to this electromagnetic radiation band, it has always been very important in the study of star light—and for life in general."

What is the \_\_\_\_\_, Alex?

- a) visible band    b) X-ray band    c) red band    d) gamma-ray band    e) big band

3. Dispersion separates in space the radiations of different wavelength (i.e., the \_\_\_\_\_) that make up a beam or propagating radiation. This allows the \_\_\_\_\_ to be analyzed. The dispersed beam is often called a \_\_\_\_\_ in a separate, but related, meaning of the word \_\_\_\_\_.

- a) range    b) electromagnetic spectrum    c) spectrum.    d) domain    e) spread

4. The dispersion of electromagnetic radiation into a spectrum can be done using a prism or a:

- a) dispenser    b) disperser.    c) dispersion grating.    d) diffraction grating.  
e) diffraction window.

5. A spectrum with no large deviations in narrow wavelength bands is a \_\_\_\_\_ spectrum and one with such deviations is a \_\_\_\_\_ spectrum. The two classes are **NOT** actually separate since a general spectrum can have both kinds of behavior. The part of a general spectrum **WITHOUT** large deviations in narrow wavelength bands is considered to be the \_\_\_\_\_ part of the spectrum.

- a) line; continuous; continuous    b) continuous; line; continuous;  
c) continuous; continuous; line    d) wavelength; continuous; wavelength  
e) line; continuous; line

6. An emission line spectrum consists of \_\_\_\_\_ against a dark background and comes from a

\_\_\_\_\_ gas. An absorption line spectrum consists of \_\_\_\_\_ against a bright background of a continuous spectrum and typically comes from a \_\_\_\_\_ gas overlying a hotter gas.

- a) dark lines; cold, dense; bright lines; hotter, dense
- b) bright lines; hot, dilute; dark lines; colder, dilute
- c) dark lines; hot, dilute; bright lines; colder, dilute
- d) dark lines; hot, dilute; dark lines; hotter, dilute
- e) bright lines; cold, dilute; bright lines; hotter, dilute

7. “Let’s play *Jeopardy!* For \$100, the answer is: He is the eponym (i.e., person after whom a thing is named) of Grotrian diagrams. A Grotrian diagram shows the energy levels of an atom, ion, or molecule in a standard format and the line transitions between the energy levels that can emit or absorb photons. It is a very abstract diagram of the atom, ion, or molecule”

Who is \_\_\_\_\_, Alex?

- a) John Venn (1834–1923)
- b) Ejnar Hertzsprung (1873–1967)
- c) Henry Norris Russell (1877–1957)
- d) Edwin Hubble (1889–1953)
- e) Walter Grotrian (1890–1954)

8. The direct observable by which stars are best empirically classified is their:

- a) emission line spectrum.
- b) absorption line spectrum.
- c) continuous spectrum.
- d) surface pressure.
- e) surface gravity.

9. The OBAFGKM spectral type classification (AKA the Harvard spectral classification) of stars is based on the line spectra (mainly absorption line spectra) of stars: O stars have a certain spectrum, B stars another, etc. The classification is empirical, but is theoretically understood to be a stellar atmosphere temperature classification. The spectral types (i.e., OBAFGKM) are ordered by decreasing stellar surface (i.e., photosphere) temperature and each **SPECTRAL TYPE** is divided into **SUBTYPES** which are numbered: the numbers in order of decreasing temperature run 0, 1, 2, 3, 4, 5, 6, 7, 8, 9. The Sun is **NOT** in the hottest, 2nd hottest, coldest, or 2nd coldest **SPECTRAL TYPE**. It is a:

- a) O5 star.
- b) B8 star.
- c) G2 star.
- d) K6 star.
- e) M3 star.

10. A \_\_\_\_\_ diagram plots logarithmic stellar luminosity versus stellar logarithmic photosphere temperature (or alternatively versus OBAFGKM spectral type or  $B - V$  color).

- a) Feynman
- b) Grotrian
- c) Hertzsprung-Russell
- d) Hubble
- e) Venn