

Name: \_\_\_\_\_

## Intro Astro Lab Prep Quiz: Lab 8: Stars

**Instructions:** There are 10 to 20 multiple-choice problems each worth 1 mark for a total of 10 to 20 marks altogether. Choose the **BEST** answer, completion, etc., and **DARKEN** fully the appropriate circle on the table provided below. Read all responses carefully. **NOTE** long detailed responses won't depend on hidden keywords: keywords in such responses are bold-faced capitalized. This is a 10 minute quiz.

### Answer Table for the Multiple-Choice Questions

	a	b	c	d	e		a	b	c	d	e
1.	O	O	O	O	O	11.	O	O	O	O	O
2.	O	O	O	O	O	12.	O	O	O	O	O
3.	O	O	O	O	O	13.	O	O	O	O	O
4.	O	O	O	O	O	14.	O	O	O	O	O
5.	O	O	O	O	O	15.	O	O	O	O	O
6.	O	O	O	O	O	16.	O	O	O	O	O
7.	O	O	O	O	O	17.	O	O	O	O	O
8.	O	O	O	O	O	18.	O	O	O	O	O
9.	O	O	O	O	O	19.	O	O	O	O	O
10.	O	O	O	O	O	20.	O	O	O	O	O

008 qmult 00100 1 1 2 easy memory: blackbody radiation defined @ prep

1. Blackbody radiation is produced by any dense body that is:

- a) at zero temperature.      b) all at one temperature.      c) at two temperatures.  
 d) at range of temperatures.      e) at infinite temperature.

**SUGGESTED ANSWER:** (b)

**Wrong answers:**

- a) This ideal limit should emit no radiation at all.

**Redaction:** Jeffery, 2013jan01

008 qmult 00120 1 4 5 easy deducto-memory: Wien's law specified @ prep

2. "Let's play *Jeopardy!* For \$100, the answer is:

$$\lambda_{max} = \frac{2897.7729 \mu\text{m K}}{T},$$

where T is Kelvin temperature. The is law gives the peak wavelength of blackbody radiation."

What is \_\_\_\_\_, Alex?

- a) the Stefan-Boltzmann law      b) Hooke's law      c) Rayleigh-Jeans law      d) Gauss's law  
 e) Wien's law

**SUGGESTED ANSWER:** (e)

**Wrong answers:**

- a) Very close in a sense.

**Redaction:** Jeffery, 2013jan01

008 qmult 00122 1 4 3 easy deducto-memory: Wien's law calculation 1 @ prep

3. From Wien's law

$$\lambda_{max} = \frac{2897.7729 \mu\text{m K}}{T}$$

(where T is Kelvin temperature), the approximate peak wavelength for a blackbody radiator at 3000 K is:

- a) 3000  $\mu\text{m}$ .    b) 3  $\mu\text{m}$ .    c) 1  $\mu\text{m}$ .    d) 1/3  $\mu\text{m}$ .    e) 1/3000  $\mu\text{m}$ .

**SUGGESTED ANSWER:** (c)

**Wrong answers:**

- a) So-so guess.

**Redaction:** Jeffery, 2013jan01

008 qmult 00150 1 1 2 easy memory: photosphere emission approximates blackbody radiation @ prep

4. The emission from a stellar photosphere approximates:

- a) white light.    b) blackbody radiation.    c) visible light.    d) LED emission.  
e) an emission line spectrum.

**SUGGESTED ANSWER:** (b)

**Wrong answers:**

- a) Only for some stars.

**Redaction:** Jeffery, 2013jan01

008 qmult 00200 1 4 5 easy deducto-memory: magnitude system defined @ prep

5. "Let's play *Jeopardy!* For \$100, the answer is: It is a system of classification the apparent (i.e., as-viewed-from Earth) or absolute brightnesses of stars and other astro-bodies. It originated with the ancient Greek astronomers who classified stars into six classes: the stars in each category judged by naked-eye visual astronomy to be of comparable apparent brightnesses. The classes in order of **DECREASING** brightness are 1st, 2nd, 3rd, 4th, 5th, and 6th magnitude. In the 19th century, it was decided to modernize this ancient classification system fixing its values to objective light flux measurements. (Flux is energy per unit time per unit area either per wavelength or integrated over some wavelength band.) The modernization was based on the discovery that the ancient magnitudes were roughly logarithmic in flux and that an **INCREASE** of 5 ancient magnitudes corresponded to roughly a factor of 100 **DECREASE** in flux. This rough result suggested the implemented prescription that an **INCREASE** of 5 magnitude corresponds to exactly a factor of 100 **DECREASE** in flux. The formula for the prescription is

$$\Delta M = -2.5 \log(F_2/F_1) ,$$

where  $\Delta M$  is the difference in magnitude between 2 astro-bodies with fluxes  $F_1$  and  $F_2$ . The negative sign makes magnitude difference increase with decreasing fraction  $F_2/F_1$ . If  $F_2/F_1 = 1/100$ , then  $\Delta M = 5$ . The inverse relationship is

$$\frac{F_2}{F_1} = 10^{-0.4 \times \Delta M} .$$

We can see now that the logarithms are actually base

$$10^{0.4} = 2.511886 \dots \approx 2.512 .$$

This means an increase in magnitude by one corresponds to a decrease in flux by a factor of  $\sim 1/2.512$ .

In the modern system, fractional magnitudes occur and the magnitudes run over the whole real number line. Very bright objects have negative magnitudes.

Actually, many people (like yours truly) think that making modern system mimic the ancient system was a stupid idea. The modern system runs the wrong way—bigger/smaller is dimmer/brighter. This leads to endless confusion. And the modern system has a logarithm base used for nothing else. One could have made the definition

$$\Delta M = \log(F_2/F_1)$$

and then bigger/lower would be brighter/dimmer and 1 magnitude would correspond to a factor of 10 in flux. That would have been so easy to understand. But no. The dead hand of the past prevails."

What is the \_\_\_\_\_, Alex?

- a) Greek system    b) magification system    c) Roman system    d) Ptolemaic system  
e) magnitude system

**SUGGESTED ANSWER:** (e)

**Wrong answers:**

- d) Ptolemy really deserves to take the hit on this one.

**Redaction:** Jeffery, 2013jan01

008 qmult 00210 1 4 5 easy deducto-memory: Ptolemy magnitude system @ prep

6. “Let’s play *Jeopardy!* For \$100, the answer is: He/She left to posterity and may have invented the ancient Greek system of 6 stellar magnitudes.”

Who is \_\_\_\_\_, Alex?

- a) Aristotle (384–322 BCE)    b) Berossos, priest of Bel Marduk (3rd century BCE)  
c) King Ptolemy I (c. 367–c. 283 BCE)    d) Cleopatra (69–30 BCE)  
e) Ptolemy (circa 100–175 CE)

**SUGGESTED ANSWER:** (e) Yours truly calls it the Ptolemaic magnitude system which may not be a common expression, but everyone knows what yours truly means.

**Wrong answers:**

- b) Berossos of Babylon (3rd century BCE) moved to the Greek island of Kos and founded a school astronomy/astrology (No-38–39).  
c) The Macedonian dynasty of Egypt had many King Ptolemys but none were astronomers to my knowledge.  
d) The gender should be the give away. Actually Cleopatra in after-legend was credited with arcane wisdom and is the apocryphal author of a work on alchemy I believe.

**Redaction:** Jeffery, 2001jan01

008 qmult 00230 1 1 2 easy memory: absolute magnitude defined @ prep

7. In order report the intrinsic brightness of stars, we define \_\_\_\_\_ to be the apparent magnitude measured at a distance of 10 parsecs. Why 10 parsecs? Maybe because its a round number that is typical for distances to nearby stars and yields \_\_\_\_\_s that are not so different from apparent magnitudes for these nearby stars.

- a) luminosity    b) absolute magnitude    c) flux    d) raw magnitude    e) watt

**SUGGESTED ANSWER:** (b)

**Wrong answers:**

- a) A related quantity

**Redaction:** Jeffery, 2013jan01

008 qmult 00300 1 4 4 easy deducto-memory: color index defined @ prep

8. “Let’s play *Jeopardy!* For \$100, the answer is: This quantity is the difference between magnitudes in two passbands for a star: the redder passband magnitude is subtracted from the bluer passband magnitude. Because of the subtraction, the distance dependence of the passband magnitudes cancels out and the quantity is a measure of the instrinsic shape of the star’s spectrum. The quantity can be used in many cases to determine the star’s surface temperature, and is often used as substitute or proxy for surface temperature in plots and discussions. The flux in the bluer passband usually increases with temperature relative to the blux in the redder passband. But this means that the quantity decreases with increasing temperature—like magnitude, the quantity increases in the wrong way leading often to confusion. The most common version of this quantity is  $B - V$ : i.e., the  $B$  passband magnitude minus the  $V$  passband magnitude.”

What is \_\_\_\_\_, Alex?

- a) absolute magnitude    b) apparent magnitude    c) luminosity    d) color index or color  
e) blueness

**SUGGESTED ANSWER:** (d)

**Wrong answers:**

- e) This would make sense if color increased with temperature.

**Redaction:** Jeffery, 2013jan01

008 qmult 00500 1 1 4 easy memory: luminosity defined @ prep

9. The total energy output per unit time could reasonably be called star power or, less reasonably, star wattage, but, in fact, is called:

- a) flux.    b) apparent magnitude.    c) absolute magntiude.    d) luminosity.    e) color.

**SUGGESTED ANSWER:** (d)

**Wrong answers:**

- a) A nonsense answer.

**Redaction:** Jeffery, 2013jan01

008 qmult 00600 1 4 2 easy deducto-memory: HR diagram defined @ prep

10. “Let’s play *Jeopardy!* For \$100, the answer is: This diagram is a plot that has logarithmic luminosity versus spectral type or color  $B - V$  or photospheric temperature for stars. The luminosity can be replaced by absolute V magnitude which is a good proxy for logarithmic luminosity. Since spectral type and color  $B - V$  increase to the right, temperature for consistency increases to the left.”

What is the \_\_\_\_\_ diagram, Alex?

- a) Bertrand Russell or BR    b) Hertzsprung-Russell or HR    c) color-color    d) star  
e) true star

**SUGGESTED ANSWER:** (b)

**Wrong answers:**

- a) Bertrand Russell (1872–1970). Great old guy.

**Redaction:** Jeffery, 2013jan01

008 qmult 00610 1 4 5 easy deducto-memory: main sequence defined @ prep

11. “Let’s play *Jeopardy!* For \$100, the answer is: This narrow band of stars on an HR diagram starts high on the left-hand side, declines rapidly, then declines slowly in middle region of the diagram, and then declines rapidly toward the right-hand side. About 90% of all stars (i.e., nuclear burning stars) in the Milky Way fall in the band and the same is roughly true of many other galaxies. The stars in the band are burning (in a nuclear sense) hydrogen to helium in their cores. The core-hydrogen-burning phase of a star’s nuclear-burning life is the longest phase and this accounts for the abundance of stars in the band.”

What is the \_\_\_\_\_, Alex?

- a) color sequence    b) giant region    c) supergiant region    d) white dwarf  
e) main sequence

**SUGGESTED ANSWER:** (e)

**Wrong answers:**

- a) As Lurch would say AAAARGH.

**Redaction:** Jeffery, 2013jan01

008 qmult 00620 1 4 3 easy deducto-memory: zero-age main sequence defined @ prep

12. “Let’s play *Jeopardy!* For \$100, the answer is: Stars on this narrow curve on an HR diagram are just at the beginning of their core-hydrogen-burning phase. The curve is roughly speaking the lower-edge of the main sequence.”

What is \_\_\_\_\_ main sequence, Alex?

- a) top    b) bottom    c) zero-age    d) beginning    e) infant

**SUGGESTED ANSWER:** (c)

**Wrong answers:**

- a) As Lurch would say AAAARGH.

**Redaction:** Jeffery, 2013jan01