

Dielectronic Recombination In Active Galactic Nuclei

D. Lukić* Z. Altun† N.R. Badnell‡
C. Brandau^{||§}M. Lestinsky¶ A. Müller^{||} D.W. Savin*
S. Schippers^{||} E.W. Schmidt^{||} M. Schnell*
F. Sprenger¶ A. Wolf¶

February 3, 2006

Abstract

XMM-Newton and Chandra observations of active galactic nuclei show a rich spectrum of X-ray absorption lines. These observations have detected a broad unresolved transition array (UTA) between $\approx 15\text{-}17$ Å. This is attributed to inner-shell photoexcitation of M-shell iron ions. Modelling these UTA features is currently limited by uncertainties in the low temperature dielectronic recombination (DR) data for M-shell iron. In order to resolve this issue and to provide reliable iron M-shell DR data for plasma modeling, we are carrying out a series of laboratory measurements using the heavy-ion Test Storage Ring at the Max-Planck-Institute for Nuclear Physics in Heidelberg, Germany. Storage rings are currently the only laboratory method capable of studying low temperature DR. We use our results to produce experimentally-derived DR rate coefficients. We are also providing our data to atomic theorists to benchmark their DR calculations. Here we will report our recent DR results for selected Fe M-shell ions. At temperatures where these ions are predicted to form in photoionized gas, we find a significant discrepancy between our experimental results and previously recommended DR rate coefficients.

This work has been supported in part by NASA, the German Federal Ministry for Education and Research, and the German Research Council.

*Columbia Astrophysics Laboratory, New York, NY, USA

†Marmara University, Istanbul, Turkey

‡University of Strathclyde, Glasgow, UK

§Gesellschaft für Schwerionenforschung, Darmstadt, Germany

¶Max-Planck-Institut für Kernphysik, Heidelberg, Germany

^{||}Justus-Liebig-Universität, Giessen, Germany