# Quantum Mechanics

## Homework 9: Time-Independent Perturbation Theory

017 qfull 00200 3 3 0 tough math: time dependent perturbation, square well

**Extra keywords:** (MEL-141:5.3), time dependent perturbation, infinite square well

1. At time t = 0, an electron of charge  $\tilde{e}$  is in the *n* eigenstate of an infinite square well with potential

$$V(x) = \begin{cases} 0, & x \in [0, a];\\ \infty & x > a. \end{cases}$$

At that time, a constant electric field  $\tilde{E}$  pointed in the positive x direction is suddenly applied. (Note the tildes on charge and electric field are to distinguish these quantities from the natural log base and energy.) **NOTE:** The 1-d infinite square-well eigenfunctions and eigen-energies are, respectively

$$\psi_n(x) = \sqrt{\frac{2}{a}} \sin\left(\frac{n\pi}{a}x\right)$$
 and  $E_n = \frac{\hbar^2 k^2}{2m} = \frac{\hbar^2}{2m} \left(\frac{\pi}{a}\right)^2 n^2$ ,

where n = 1, 2, 3, ... The sinusoidal eigenfunctions can be expressed as exponentials: let  $z = \pi x/a$ , and then

$$\sin(nz) = \frac{e^{inz} - e^{-inz}}{2i}$$

- a) Use 1st order time-dependent perturbation theory to calculate the transition probabilities to all **OTHER** states m as a function of time. You should evaluate the matrix elements as explicitly: this is where all the work is naturally.
- b) How do the transition probabilities vary with the energy separation between states n and m?
- c) Now what is the 1st order probability of staying in the same state n?

### SUGGESTED ANSWER:

a) In this case the perturbation potential is given by

$$H^{(1)} = (-\tilde{e})(-\tilde{E})x = \tilde{e}\tilde{E}x .$$
<sup>(1)</sup>

The first order perturbation expression for the coefficient  $a_m \ (m \neq n)$  is then

$$a_m = \frac{1}{i\hbar} \int_0^t e^{i\omega_{mn}t'} H_{mn}^{(1)}(t') dt'$$
  
=  $\frac{1}{i\hbar} e^{i\omega_{mn}t/2} \frac{\sin(\omega_{mn}t/2)}{(\omega_{mn}/2)} \tilde{e}\tilde{E}\langle\psi_m|x|\psi_n\rangle$ , (2)

where

$$\omega_{mn} \equiv \frac{E_m - E_n}{\hbar} \ . \tag{3}$$

The transition probability is

$$P_{n \text{ to } m}(t) = |a_m|^2 = \frac{1}{\hbar^2} \frac{\sin^2(\omega_{mn}t/2)}{(\omega_{mn}/2)^2} |\tilde{e}\tilde{E}|^2 |\langle \psi_m | x | \psi_n \rangle|^2 .$$
(4)

At time zero, the transition probabilities grow everywhere as  $t^2$  as can be seen by a 1st order expansion of the sine function about zero. As time passes, the probabilities begin to oscillate in time. Non-monotonic behavior (which we can call the beginning of oscillation) sets in for state m when

$$\frac{\pi}{2} < \frac{\omega_{mn}t}{2} \tag{5}$$

holds for that state. The probability would continue to grow as  $t^2$  only for states degenerate with state n: i.e., for cases with  $\omega_{mn} = 0$ . But there are no degenerate states for 1-dimensional infinite square well. One must recall that 1st order theory can not be valid for long times: i.e., for  $P_{n \text{ to } m}(t)$  growing significantly close to 1.

The remaining problem is just to evaluate and simplify the expression for the matrix element:

$$\langle \psi_m | x | \psi_n \rangle = \frac{2}{a} \int_0^a \sin\left(\frac{m\pi}{a}x\right) x \sin\left(\frac{n\pi}{a}x\right) dx$$

$$= \frac{2}{a} \frac{1}{(-4)} \left(\frac{a}{\pi}\right)^2 \int_0^\pi \left(e^{imz} - e^{-imz}\right) z \left(e^{inz} - e^{-inz}\right) dz$$

$$= \frac{2a}{\pi^2} \frac{1}{(-4)} \int_0^\pi z \left(e^{i(m+n)z} + e^{-i(m+n)z} - e^{i(m-n)z} - e^{-i(m-n)z}\right) dz , \qquad (6)$$

where we have used the transformation  $z = \pi x/a$ . Here we note that

$$\int_0^{\pi} z e^{i\ell z} \, dz = \frac{\pi}{i\ell} (-1)^\ell + \frac{(-1)^\ell - 1}{\ell^2} \,, \tag{7}$$

where  $\ell$  is an integer. The first terms from right-hand side of equation (7) that appear in integral of equation (6) all cancel out in fact. The second terms from right-hand side of equation (7) that appear in integral of equation (6) yield the factor

$$2\frac{[(-1)^{m-n}-1]}{(m+n)^2} + 2\frac{[(-1)^{m-n}-1]}{(m-n)^2} = \frac{-8mn}{(m^2-n^2)^2}[(-1)^{m-n}-1],$$
(8)

where we have use the fact that if m + n is even then m - n is even too, and so  $(-1)^{m+n} = (-1)^{m-n}$ . The final results are

$$\langle \psi_m | x | \psi_n \rangle = \frac{8a}{\pi^2} \frac{mn}{(m^2 - n^2)^2} \left\{ \frac{[(-1)^{m-n} - 1]}{2} \right\}$$
(9)

and

$$\left|\langle\psi_m|x|\psi_n\rangle\right|^2 = \frac{64a^2}{\pi^4} \frac{m^2 n^2}{(m^2 - n^2)^4} \left\{\frac{[(-1)^{m-n} - 1]}{2}\right\}^2 \,. \tag{10}$$

We can see from the matrix element that all even transitions (i.e., those with m - n even) vanish in 1st order theory.

b) How does energy separation of states m and n affect the matrix elements? Recall from equation  $E_n$  is proportional  $n^2$  and is always greater than zero. Consider the function

$$f(r) = \frac{rb}{(r-b)^4} .$$
 (11)

The derivative is

$$\frac{df(r)}{dr} = -\frac{b(b+3r)}{(r-b)^5} \ . \tag{12}$$

Evidently, the function f(r) for r > 0 and b > 0 decreases with r for r > b and increases with r for r < b. Thus f(r) decreases always as |r - b increases. In our case  $r = m^2$  and  $b = n^2$ . We can see that the squared modulus of the odd matrix element decreases with increasing energy separation. Thus the transition probabilities decrease with increasing energy separation between states n and m.

c) The first order amplitude for staying in the same state is

$$a_m = \frac{1}{i\hbar} \int_0^t H_{nn}^{(1)}(t') dt'$$
  
=  $\frac{1}{i\hbar} t \tilde{e} \tilde{E} \langle \psi_n | x | \psi_n \rangle$  (13)

Now

$$\langle \psi_n | x | \psi_n \rangle = \frac{2}{a} \int_0^a x \sin^2 \left(\frac{n\pi}{a} x\right) dx = \frac{2}{a} \left(\frac{a}{n\pi}\right)^2 \int_0^{n\pi} z \sin^2(z) dz$$

$$= \frac{2}{a} \left(\frac{a}{n\pi}\right)^2 \left\{ \frac{z}{2} \left[ z - \frac{\sin(2z)}{2} \right] \Big|_0^{n\pi} - \int_0^{n\pi} \frac{1}{2} \left[ z - \frac{\sin(2z)}{2} \right] dz \right\}$$

$$= \frac{2}{a} \left(\frac{a}{n\pi}\right)^2 \left\{ \frac{(n\pi)^2}{2} - \frac{1}{2} \left[ \frac{z^2}{2} + \frac{\cos(2z)}{4} \right] \Big|_0^{n\pi} \right\}$$

$$= \frac{2}{a} \left(\frac{a}{n\pi}\right)^2 \frac{(n\pi)^2}{4}$$

$$= \frac{a}{2}$$

$$(14)$$

Then we find

$$P(t) = \frac{t^2}{\hbar^2} |\tilde{e}\tilde{E}|^2 \frac{a^2}{4} .$$
(15)

The 1st order probability of no transition increases quadratically with time. Note a 1st order calculation for no transition is probably not that great. One should probably do a 2nd order calculation.

Redaction: Jeffery, 2001jan01

024 qmult 01600 1 4 4 easy deducto-memory: Einstein stimulated em.

2. "Let's play *Jeopardy*! For \$100, the answer is: An effect discovered by Einstein by means of a thermodynamic equilibrium detailed balance argument."

| a) spontaneous emission | b) special relativity   | c) the photoelectric effect |
|-------------------------|-------------------------|-----------------------------|
| d) stimulated emission  | e) spontaneous omission |                             |

# SUGGESTED ANSWER: (d)

#### Wrong answers:

e) I'm subject to this effect myself.

Redaction: Jeffery, 2001jan01

024 qmult 01900 1 1 2 easy memory: electric dipole transitions

3. Typically, strong atomic and molecular transitions are \_\_\_\_\_\_ transitions.

a) electric quadrupole b) electric dipole c) magnetic dipole d) electric monopole e) magnetic metropole

## SUGGESTED ANSWER: (b)

## Wrong Answers:

e) Metropole is not in my dictionary, but I think it is a word meaning metropolis. If I recall correctly, Wallace Stevens uses metropole in cute poem.

Redaction: Jeffery, 2001jan01

024 qfull 00300 2 5 0 moderate thinking: classical EM scattering

Extra keywords: reference Mi-83

4. Say we had a classical simple harmonic oscillator (SHO) consisting of a particle with mass m and charge e and a restoring force  $m\omega_0^2$  where  $\omega_0$  is the simple harmonic oscillator frequency. This SHO is subject to driving force caused by traveling electromagnetic field (i.e., light):

$$\vec{F}_{\text{drive}} = e\vec{E}_0 e^{i\omega t}$$

where  $\vec{E}_0$  is the amplitude,  $\omega$  is the driving frequency, and we have used the complex exponential form for mathematical convenience: the real part of this force is the real force. The magnetic force can be neglected for non-relativistic velocities. The Lorentz force is

$$\vec{F} = e\left(\vec{E} + \frac{\vec{v}}{c} \times \vec{B}\right)$$

(Ja-238) and  $\vec{E}$  and  $\vec{B}$  are comparable in size for electromagnetic radiation, and so the magnetic force is of order v/c smaller than the electric force. (See also MEL-130.) An oscillating charge is an accelerating charge and will radiate electromagnetic radiation. The power radiated classically is

$$P = \frac{2e^2a^2}{3c^3} \; ,$$

where  $\vec{a}$  is the charge acceleration. This radiation causes an effective damping force given approximately by

$$\vec{F}_{\text{damp}} = -m\gamma \vec{v} ,$$

where

$$\gamma = \frac{2e^2\omega_0^2}{3mc^3} \ .$$

The full classical equation of motion of the particle is

$$m\vec{a} = -m\omega_0^2\vec{r} + e\vec{E}_0e^{i\omega t} - m\gamma\vec{v}$$

- a) Solve the equation of motion for  $\vec{r}$  and  $\vec{a}$ . **HINTS:** The old trial solution approach works. Don't forget to take the real parts although no need to work out the real part explicitly: i.e., Re[solution] is good enough for the moment.
- b) Now solve for the time average of the power radiated by the particle. **HINT:** You will need the explicit real acceleration now.
- c) The average power radiated must equal the average power absorbed. Let's say that the particle is in radiation flux from a single direction with specific intensity

$$I_0 = \frac{cE_0^2}{8\pi}$$

(Mi-9), where time averaging is assumed *como usual*. The power absorbed from this flux is  $\sigma(\omega)I_0$ , where  $\sigma(\omega)$  is the cross section for energy removed. Solve for  $\sigma(\omega)$  and then find show that it can be approximated by a Lorentzian function of  $\omega$  with a coefficient  $\pi e^2/(mc)$ . **HINT:** It is convenient to absorb some of the annoying constants into another factor of  $\gamma$ .

d) Now rewrite the cross section as a function of  $\nu = \omega/(2\pi)$  (i.e., the ordinary frequency) and then integrate over  $\nu$  to get the frequency integrated cross section  $\sigma_{\nu \text{ int}}$  of the system. What is the remarkable thing about  $\sigma_{\nu \text{ int}}$ ? Think about how it relates to the system from which we derived it. Evaluate this frequency integrated cross section for an electron. **HINT**: The following constants might be useful

$$\alpha = \frac{e^2}{\hbar c} = \frac{1}{137.036}$$
,  $\hbar = 1.05457 \times 10^{-27} \,\mathrm{erg\,s}$ , and  $m_e = 9.10939 \times 10^{-28} \,\mathrm{g}$ .

### SUGGESTED ANSWER:

a) Let  $\vec{r} = \vec{r_0}e^{i\omega t}$  and substitute this into the equation of motion and cancel out the exponential factor:

$$-\omega^2 m \vec{r}_0 = -m\omega_0^2 \vec{r}_0 + e\vec{E}_0 - i\omega m\gamma \vec{r}_0 \; .$$

Well we find

$$ec{r_0} = rac{(e/m)ec{E_0}}{\omega_0^2 - \omega^2 + i\gamma\omega} \; ,$$

and so

$$\vec{r} = \operatorname{Re}\left[\frac{(e/m)\vec{E}_0e^{i\omega t}}{\omega_0^2 - \omega^2 + i\gamma\omega}\right]$$

and

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$$\vec{a} = \operatorname{Re}\left[\frac{-\omega^2(e/m)\vec{E}_0e^{i\omega t}}{\omega_0^2 - \omega^2 + i\gamma\omega}\right]$$

b) Behold:

$$\vec{a} = -\omega^2 (e/m) \vec{E}_0 \frac{(\omega_0^2 - \omega^2) \cos(\omega t) + \gamma \omega \sin(\omega t)}{(\omega_0^2 - \omega^2)^2 + \gamma^2 \omega^2}$$

The time average of the square of the magnitude of  $\vec{a}$  is

$$\langle a^2 \rangle = \left[ \omega^2 (e/m) \vec{E}_0 \right]^2 \frac{1}{2} \frac{(\omega_0^2 - \omega^2)^2 + \gamma^2 \omega^2}{[(\omega_0^2 - \omega^2)^2 + \gamma^2 \omega^2]^2} = \left[ \omega^2 (e/m) \vec{E}_0 \right]^2 \frac{1}{2} \frac{1}{(\omega_0^2 - \omega^2)^2 + \gamma^2 \omega^2} ,$$

where we have used the facts that

$$\frac{1}{2\pi} \int_0^{2\pi} \sin^2(x) \, dx = \frac{1}{2\pi} \int_0^{2\pi} \cos^2(x) \, dx = \frac{1}{2} \qquad \text{and} \qquad \frac{1}{2\pi} \int_0^{2\pi} \sin(x) \cos(x) \, dx = 0 \; .$$

The time-averaged power is then

$$\langle P \rangle = \frac{e^4 \omega^2 E_0^2}{3m^2 c^3} \frac{1}{(\omega_0^2 - \omega^2)^2 + \gamma^2 \omega^2} ,$$

c) Behold:

$$\sigma(\omega) = \frac{\langle P \rangle}{I_0} = \frac{8\pi e^4 \omega^4}{3m^2 c^4} \frac{1}{(\omega_0^2 - \omega^2)^2 + \gamma^2 \omega^2} \ . \label{eq:sigma_eq}$$

Given the denominator,  $\sigma(\omega)$  may well be sharply peaked about the SHO frequency  $\omega_0$ . Thus only the strongest variation on  $\omega$  needs to be retained. One can approximate  $\omega$  elsewhere by  $\omega_0$ . Well the big trick, and I admit to never having thought of it myself, is to make the following approximation

$$\omega_0^2 - \omega^2 = (\omega_0 + \omega)(\omega_0 - \omega) \approx 2\omega_0(\omega_0 - \omega) .$$

One then approximates the  $\omega^4$  by  $\omega_0^4$ , divides numerator and denomitor by  $4\omega_0^2$ , and absorbs some constants into another  $\gamma$  factor to get

$$\sigma(\omega) = \frac{\pi e^2}{mc} \frac{\gamma}{(\omega_0 - \omega)^2 + (\gamma/2)^2} \; .$$

d) Well

$$\sigma(\nu) = \frac{\pi e^2}{mc} \frac{1}{\pi} \frac{\gamma/(4\pi)}{(\nu_0 - \nu)^2 + (\gamma/4\pi)^2}$$

The normalized Lorentzian is

$$L(x) = \frac{1}{\pi} \frac{\Gamma/2}{(x - \bar{x})^2 + (\Gamma/2)^2}$$

where  $\Gamma$  is the FWHM (i.e., full-width at half maximum) and  $\bar{x}$  is the mean value of the distribution. It follows that

$$\sigma_{\nu \text{ int}} = \frac{\pi e^2}{mc}$$

The remarkable thing about this frequency integrated cross section is that it is independent of the restoring force of the simple harmonic oscillator. It is sort of a universal result. Now microscopic systems are not classical nor often exactly simple harmonic oscillators, and thus one does not expect  $\sigma_{\nu \text{ int}}$  to apply. However,  $\sigma_{\nu \text{ int}}$  evaluated for an electron is frequently used as a unit of frequency integrated cross sections. Atomic frequency integrated cross sections can be written out

$$\sigma_{\nu \text{ int atomic}} = \frac{\pi e^2}{mc} f_{ij} ,$$

where  $f_{ij}$  is a dimensionaless quantity called the oscillator strength of the transition from lower level *i* to upper level *j*. Only for the strongest atomic transitions does  $_{ij}$  approach unity (Mi-84).

For an electron, we find

$$\sigma_{\nu \text{ int}} = \frac{\pi e^2}{mc} = \frac{\pi \hbar \alpha}{m} = 0.02654 \,\mathrm{cm}^2/\mathrm{s} \;.$$

We note that the units are area per time. The per time is there because we integrated over frequency.

Fortran Code

```
print*
pi=acos(-1.)
finestr=1./137.036
hbar=1.05457e-27
emass=9.10939e-28
oscon=(pi*hbar*finestr/emass)
print*,'The classical frequency integrated cross section is'
print*,'oscon=',oscon ! 0.0265400279
```

Redaction: Jeffery, 2001jan01